

Summary Session II :
Observations, laboratory measurements, modelling

Observations & Diagnostics : PS, RHIC, PSR, APS, SPS

Some “old” machines show new effects

Modelling : e-cloud build-up, secondary electron yield, heat load

Very extensive compilation and modelling of secondary emission data (N.H., M.F., I.C.)

Issues : generation and survival of electrons during bunch gaps

Thresholds (multipactor electron signal, vacuum pressure, beam stability,...)

Beam conditioning (scrubbing) of surface

Comparison of observations with simulations

Discussion : micro wave as diagnostic tool for conditioning or as a remedy

Use of NEG-film, TiN coating

Survival time of electrons (PSR)

Future steps and issues for LHC

Obtain good data on ‘observables’ : heat load in SPS , WAMPAC, COLDEX, strip detector
Electron cloud stripes in SPS monitor -> urgent decision on LHC beam screen slots
Reproduce thresholds (vacuum pressure, electron signal, beam stability)

Scrubbing effect :

Detailed scenarios for LHC (bunch intensity & spacing, beam energy < 2 TeV no S.R.)

How to achieve max below 1.3

Discrepancies of real machines (SPS, B-factories) with lab data? What is their seec?

Dominated by photo-electrons rather than secondary electrons?

-> SPS beam time will be vital

Seec measurements : in situ & Lab. Note : technical surfaces & practical systems

Ions for LHC (RHIC - LEIR similarities?)

Heat load <-> scrubbing

Heat load and dose rate are related

$$D(C/s/mm^2) = \frac{P(W/m)}{F(mm^2/m) \langle E \rangle}$$

For LHC beam screen $F \sim 5 \cdot 10^4 \text{ mm}^2/\text{m}$

$P(\text{cryo-budget}) = 0.5 \text{ W/m}$ and for the cloud $\langle E \rangle \sim 100 \text{ eV}$

N. Hilleret et.al. find 10^{-2} C/mm^2 for a well scrubbed Cu surface <1.3 for max

→ should take about 30 hours to accumulate this dose

→ If heat load is limiting us in the LHC, scrubbing should go fast, otherwise no problem

This also applies to vacuum scrubbing of the surface.

SPS : low duty cycle; takes longer.